

**CosmoPhi: A New Code for
Hydrodynamic Cosmology
Simulation by means
Intel Xeon Phi Supercomputers**

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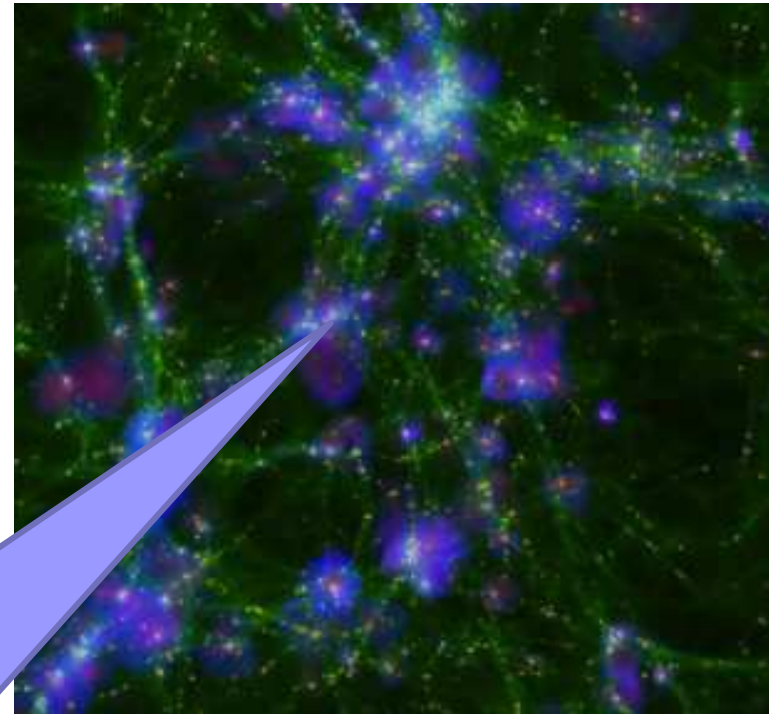
**7 July 2015
Aix-en-Provence, France**

The motivation

«The movement of galaxies in dense clusters turns the collisions of galaxies into an important evolutionary factor»

Professor

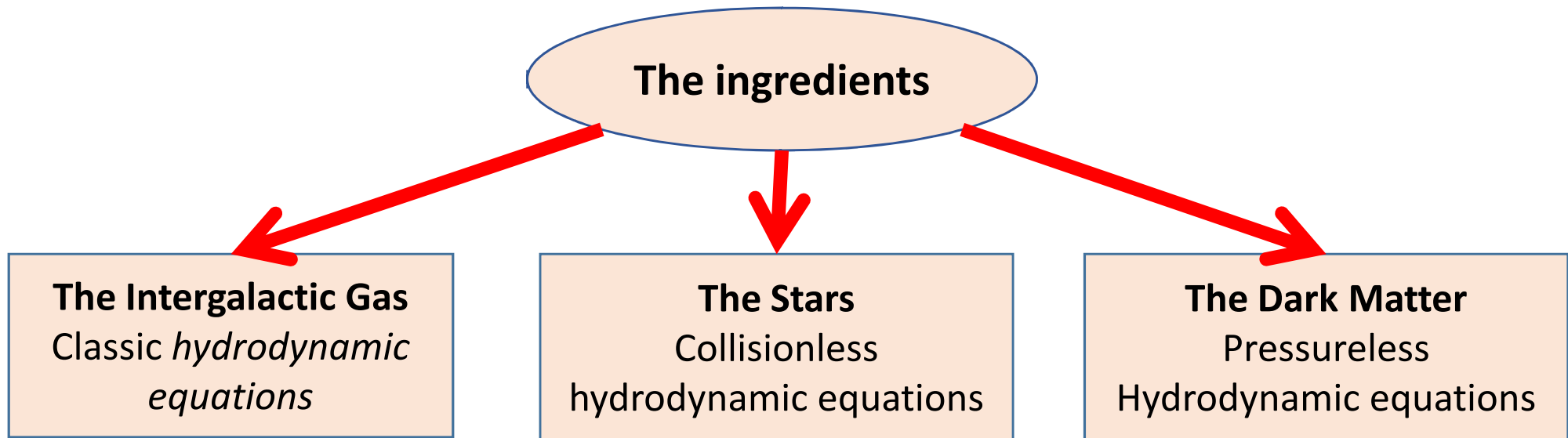
Dr. Alexander Tutukov



The challenges:

1. The mathematical model
2. The numerical solver
3. The efficiency implementation

The physics model



The subgrid physics model

- Gravity potential
- Cooling/Heating terms
- Magnetic field
- Nonequilibrium chemistry
- Star formation
- Supernovae feedback

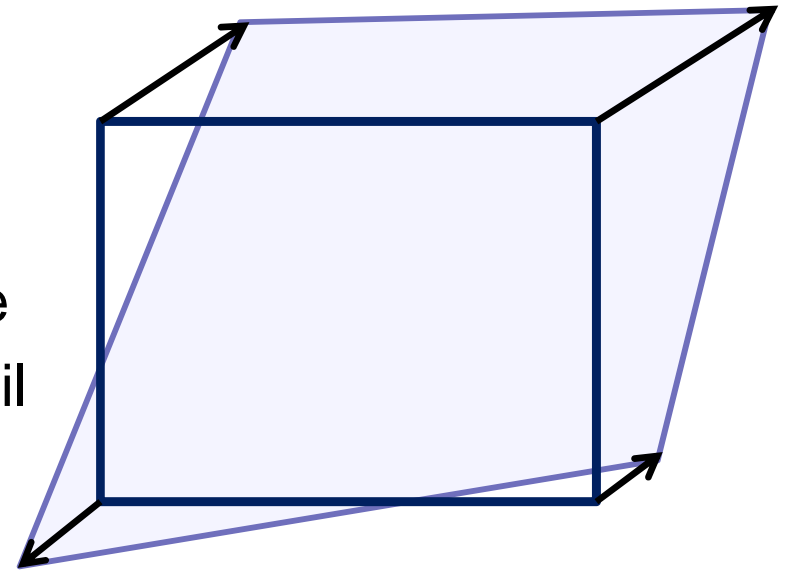
Main features:

1. Unified numerical solver
2. Unified parallel algorithm
3. Potential infinity scalability

The numerical method

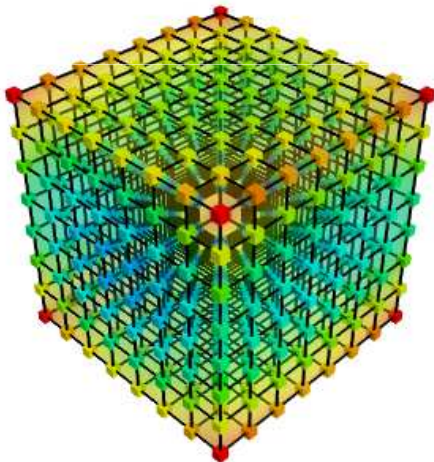
The numerical method

- Operator splitting approach
- Godunov method
- Roe solver adaptation for vacuum interface
- Piecewise-parabolic method on local stencil
- Fast Fourier Transform technique
- Dual energy formalism

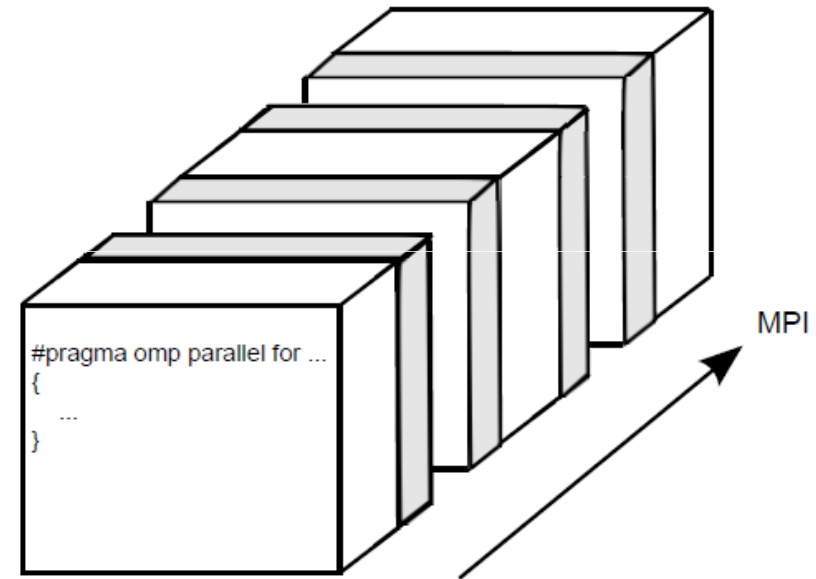
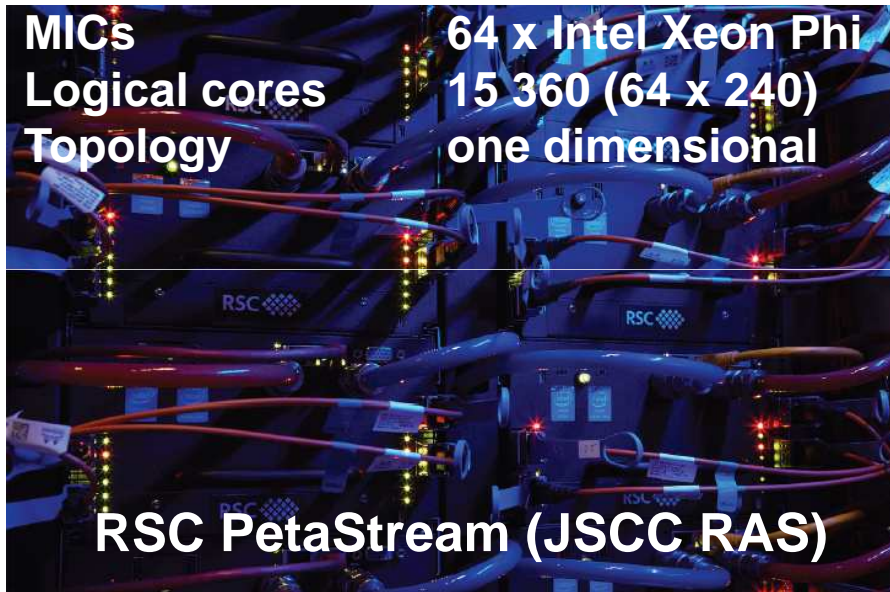


The features of numerical method

- High-order accurate method
- The limiter free solver
- Galilean invariant numerical solution
- Guarantee of non-decrease entropy
- Simplicity and Scalability of implementation
- Extensionability on more hyperbolic models



The parallel implementation on RSC PetaStream

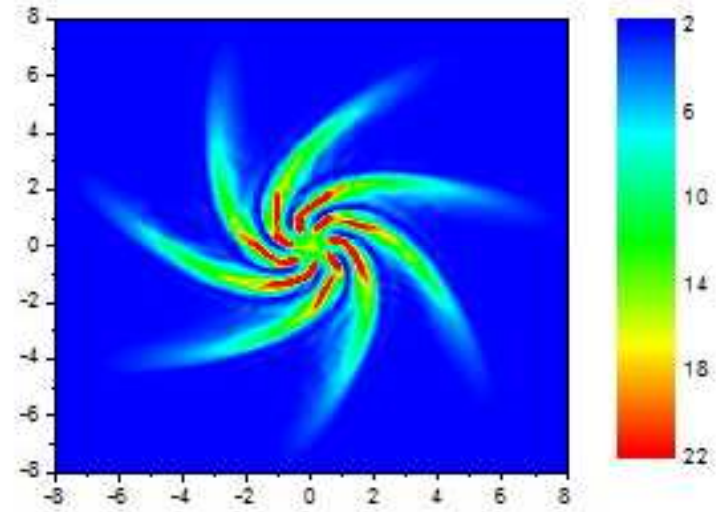
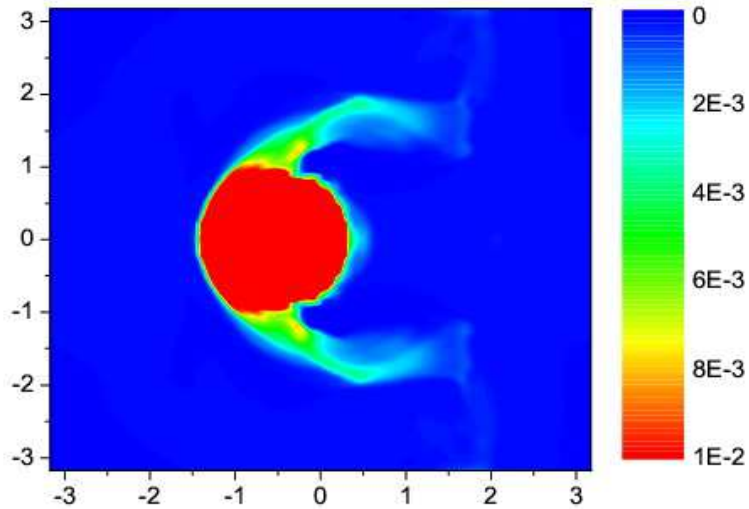


1. The speed-up factors of 134 on 260 logical cores
2. The efficiency of 92% on 64 MICs (or on 15 360 cores)
3. The 29 GFLOPS of scalar performance (or 40% from peak)

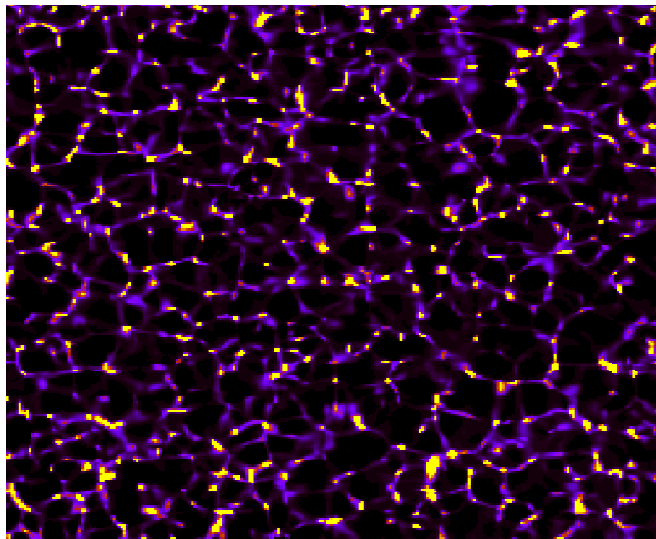
The simulation of behavior of parallel implementation on 983 040 cores by means AGNES* system the 80% efficiency was shown

*Podkorytov et al., LNCS 2010

The interacting & evolution galaxy



Cosmology simulation





The collaborators

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Thank you for attention!

Welcome to my poster